

# Airy's isostatic model: a proposal for a realistic case

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**Abstract** This work derives algebraic expressions for Airy-type isostatic equilibrium and disequilibrium of lithospheric blocks within asthenosphere, considering a more realistic case of continuously varying densities of the lithosphere along mutually perpendicular directions. Isostatic (dis)equilibrium must be linked to the porosity and density of the rock matrix, and to pore fluid density since these factors govern the bulk density of wet sediments/porous rocks. Two combinations of exponential and linear increase of density with depth are demonstrated. Isostatic disequilibrium would speed up sinking/uplift of the lithospheric blocks. Given any other (empirical) equations for density variation in three perpendicular directions, isostatic (dis)equilibrium such as Eqs. 14, 17, 18, 28, and 29 can be theorized as well.

**Keywords** Isostasy · Density-depth relation · Porosity · Pore fluid density · Rock matrix density

## Introduction

“Earth’s surface processes, namely erosion, transport and sedimentation, tend to counteract... positive and negative vertical movements.”—G.V. Bosch et al. (2016)

The leading isostasy models, such as that by Airy, consider lithospheric/crustal blocks with constant densities

floating in the asthenosphere/mantle (Turcotte and Schubert 2002). However, even the same rock type usually gets denser as depth increases. Heterogeneity in density in lithosphere is common (Singh et al. 2015), and classically Airy model of isostasy has been modified by considering pressure exerted by individual layers of rocks of different densities: wide isostatic balance exercise for rifted basins in Turcotte and Schubert (2002). From a different perspective, Airy model was modified by Meiniez by considering a distributed compensation region (Vanicek and Krakiwsky 1986). Litinsky (1989) formulated “effective density” for layered rocks that involve densities and thicknesses of each layers. Cordell (1973) suggested that density contrast decreases exponentially with depth, with decrement ranging from 0.3 to 1.5 c.c. km<sup>-1</sup>. Density usually increases (Goteti et al. 2012) but rarely can even decrease (Ebbing et al. 2007 and its review) exponentially with depth. Carlson and Herrick (1990) stated an average density gradient of 0.32 Mg m<sup>-3</sup> km<sup>-1</sup>. Reid (1987) pointed out a linear increase in density with depth from oceanic crust. Such a linear relation has also been adopted/modeled for the entire lithosphere (e.g., Motavalli-Anbaran et al. 2013; Kumar et al. 2014; Xu et al. 2016), and sometimes specifically for the mantle (Clark and Ringwood 1964) that can depend on its thermal structure (Xu et al. 2016). Zhang and Chen (1992) referred such a relation from deep crust, which Hofmeister and Criss (2015) mentioned just as a possibility. Besides, density decreases in overpressure zones in marine and shallow sediments (Bowers 2001). Even soil layers may exhibit a similar density-depth relation (Román-Sánchez et al. 2017). The linear density-depth relation (e.g., Martinec and Fullea 2015) has also been incorporated in few tectonic models (Pearse and Bailey 2007). The linear relation could be ultimately guided by the temperature structure with depth (Nemčok 2016).

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Porous sedimentary rocks (clastics and carbonates; McCulloh 1967; Maxant 1980) might be filled up with pore fluids, usually water, and this modifies the depth integrated “mean density” (Goteti et al. 2012) of the lithospheric block. Layered sedimentary rocks undergoing deeper burial can compact more and show increase in density (Maden et al. 2015). A gradual variation in density can also be induced by metamorphism (review in Zhou 2009). Tenzer et al. (2012) derived for the entire Earth almost a linear density gradient of  $13 \pm 2 \text{ kg m}^{-3} \text{ km}^{-1}$  within the upper mantle, 6–58 km below the continental crust. A few geodynamic models have considered density variation in all the three directions (e.g., Zhou 2009). Goteti et al. (2012) used exponential depth-density relation of sediments in analytical models. Not all rocks increase density with depth, e.g., evaporate/rock salt where in fact the density falls slightly due to thermal expansion (Romer and Neugebauer 1991). Sedimentary facies variation in 2D is quite common in marine to non-marine transition zones where density of one rock type is expected to alter progressively into the other. The South Caspian Basin hosts the thickest (28 km: Knapp et al. 2007) known sediment pile on the Earth where linear density variation with depth persist up to a long distance/depth (Motavalli-Anbaran et al. 2013).

This work develops models of Airy type isostatic (dis)equilibrium of blocks with continuously varying densities in orthogonal directions. If someone picks up density of the surface rock as a representative of that of the entire lithosphere, she would be proceeding with unrealistic magnitude that is not representative of the lithosphere. The model presented here overcomes this problem. However, abrupt facies variation (Chaudhuri et al. 1987) along horizontal/vertical directions and depth-wise transition from crust to mantle (Fig. 3 of Xu et al. 2016), leading to sharp jump in density, may not be possible to fit with suitable empirical equations, and the model presented here does not work in those cases.

**Model** Consider a sedimentary lithospheric block floating on asthenosphere (Fig. 1). The block is porous and consists of a single phase fluid in the pore space (Rieke and Chilingarian 1974).

Here

$$\rho_{bw} = \rho_m - (\rho_m - \rho_f)\varnothing \tag{1}$$

$\rho_{bw}$ : wet bulk density of sediments/rock,  $\rho_m$ : matrix/mineral/grain density,  $\rho_f$ : fluid density, and  $\varnothing$ : fractional porosity.

If porosity falls exponentially with depth (Athy 1930):

$$\varnothing_z = \varnothing_0 e^{-bz} \tag{2}$$

$\varnothing_z$ : porosity at depth  $z$ ,  $\varnothing_0$ : porosity at surface (i.e., at  $z = 0$ ),  $e$ : exponential series, and  $b$ : a constant. Note  $b^{-1} = \lambda$ : compaction constant.

From Eqs. 1 and 2,

$$\rho_{bwz} = \rho_m - (\rho_m - \rho_f)\varnothing_0 e^{-bz} \tag{3}$$

This means, at surface, or

$$z = 0, \rho_{bw0} = \rho_m - (\rho_m - \rho_f)\varnothing_0 \tag{4}$$

**Case I** This case shows density increases with depth ( $Z$  direction) exponentially. In two perpendicular horizontal directions  $X$  and  $Y$ , densities of two types of sediments/rocks vary linearly, which could be the case of a sedimentary facies variation. Lateral density variation has been inferred to affect rocks as deep as 20 km (Wu and Mereu 1990).

$$\rho(x, 0, 0) = \rho_m - (\rho_m - \rho_f)\varnothing_0 + k_x x \tag{5}$$

$$\rho(0, y, 0) = \rho_m - (\rho_m - \rho_f)\varnothing_0 + k_y y \tag{6}$$

$$\rho(0, 0, z) = \rho_m - (\rho_m - \rho_f)\varnothing_0 e^{-bz} \tag{7}$$

Here, density gradients  $k_x, k_y \neq 0$ . Obviously for increase in density away from the origin (point O in Fig. 1),  $k_x, k_y, b > 0$ , and for the decreasing case,  $k_x, k_y, b < 0$ . Density at the origin is  $\rho_{bw} = \{\rho_m - (\rho_m - \rho_f)\varnothing_0\}$ . From Eqs. 5 to 8, density at any point  $(x, y, z)$  within the block:

$$\rho(x, y, z) = \rho_m - (\rho_m - \rho_f)\varnothing_0 e^{-bz} + k_x x + k_y y \tag{8}$$

Note that  $\rho(x, 0, 0)$ ,  $\rho(0, y, 0)$ , and  $\rho(0, 0, z)$  obtained from Eq. 8 match with Eqs. 5, 6, and 7, respectively.

Mass of the lithospheric block:

$$M = \int_0^{x_1} \int_0^{y_1} \int_0^{z_1} \rho_{bw}(x, y, z) dx dy dz \tag{9}$$

Equation 9 is a standard step adopted in finding mass of a body with variable density (e.g., Das and Mukherjee 1996).

Or

$$M = \int_0^{x_1} \int_0^{y_1} \int_0^{z_1} \{\rho_m - (\rho_m - \rho_f)\varnothing_0 e^{-bz} + k_x x + k_y y\} dx dy dz \tag{10}$$

$$\text{Simplifying, } M = \prod q_1 [\rho_m + 0.5\{k_x x_1 + k_y y_1\}] + (\rho_m - \rho_f)\varnothing_0 x_1 y_1 b^{-1} (e^{-bz_1} - 1) \tag{11}$$

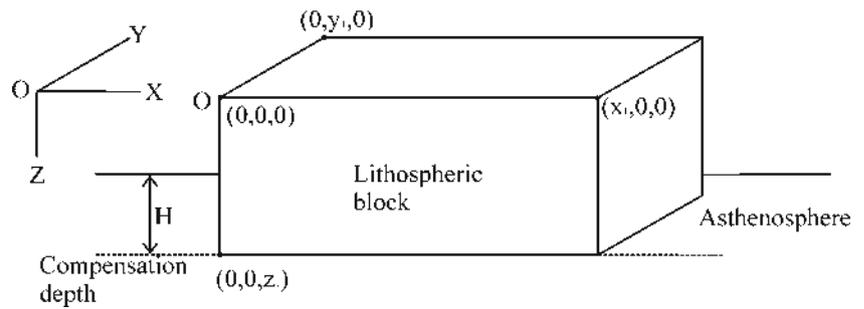
Here,  $q = (x, y, z)$ . Thus,  $\prod q_1 = x_1 y_1 z_1$

Volume of the block is  $v = \prod q_1$ .

Therefore, the mean density ( $\rho_e$ ) of the block:

$$\rho_e = (\prod q_1^{-1} M) \tag{12}$$

**Fig. 1** A lithospheric block floating in asthenosphere. The compensation depth lies at the bottom surface of the block, as per the Airy model



which is

$$\rho_e = [\rho_m + 0.5\{k_x x_1 + k_y y_1\}] + (\rho_m - \rho_f)\varnothing_0 z_1^{-1} b^{-1} (e^{-bz_1} - 1) \tag{13}$$

For isostatic equilibrium:

$$\rho_a H = z_1 \rho_e \tag{14}$$

Equation (14) is a standard way of pressure balance at the compensation depth (bottom margin of the lithosphere), as stated in pp. 74 of Turcotte and Schubert (2002).

For a constant rate of sedimentation  $dz/dt = h_s > 0$ , pressure difference per unit time driving the lithospheric block down:

$$d(\Delta P)/dt = g \{h_s \rho_e - H\rho_a\} \tag{15}$$

$g$  is the acceleration due to gravity. The rate of change of the force driving the block to sink

$$d(F_{down})/dt = d(\Delta P)/dt \times x_1 y_1 = g x_1 y_1 \{h_s \rho_e - H\rho_a\} \tag{16}$$

“ $x_1 y_1$ ” being the horizontal basal area of the block on which the force works.

Dividing this by the mass of the block as in Eq. 11, the rate of increase of acceleration of the sinking block:

$$d(f_{down})/dt = g x_1 y_1 \{h_s \rho_e - H\rho_a\} \{[\prod q_1 [\rho_m + 0.5\{k_x x_1 + k_y y_1\}] + (\rho_m - \rho_f)\varnothing_0 x_1 y_1 b^{-1} (e^{-bz_1} - 1)]\}^{-1} \tag{17}$$

Consider contrarily that the lithospheric block is under a constant rate of erosion that decreases its height at a rate  $h_e = -dz/dt$ . Then

$$d(f_{up})/dt = g x_1 y_1 \{H\rho_a - h_s \rho_e\} \{[\prod q_1 [\rho_m + 0.5\{k_x x_1 + k_y y_1\}] + (\rho_m - \rho_f)\varnothing_0 x_1 y_1 b^{-1} (e^{-bz_1} - 1)]\}^{-1} \tag{18}$$

For a (nearly) constant porosity with depth (Sugai et al. 1994 as just a single example),  $b = 0$ . Then, Eqs. 11, 13, 17, and 18 simplify significantly and become independent of density of the pore fluid ( $\rho_f$ ).

**Case II** This case shows linear increase of density with depth ( $Z$  direction) and also the same pattern in two perpendicular horizontal directions  $X$  and  $Y$ . Siltstones, sandstones, dolostones, and shale from few locations have shown almost linear increase in density with depth up to ~4 km, whereas for overpressured sediments, the relation is strongly non-linear (Fig. 2.7a of Telford et al. 1990). Further, compaction of argillaceous sediments (shale density) increases exponentially up to ~3–4 km depth (Rieke and Chilingarian 1974).

$$\rho(q) = \rho_m - (\rho_m - \rho_f)\varnothing_0 + k_q q \quad (q = x, y, z) \tag{19}$$

Therefore,  $\rho(x, y, z) = \rho_m - (\rho_m - \rho_f)\varnothing_0 + \sum k_q q \quad (q = x, y, z)$  (20)

Here, too, the density at the origin (0, 0, 0) is  $\rho_m - (\rho_m - \rho_f)\varnothing_0$ .

Using the right-hand side expression in Eq. 9, doing triple integration applying limits (0,  $x_1$ ), (0,  $y_1$ ), and (0,  $z_1$ ):

$$\text{Mass of the block } M = \prod q_1 \{ \rho_m - (\rho_m - \rho_f)\varnothing_0 + 0.5 \sum k_q q_1 \} \quad (q = x, y, z) \tag{21}$$

Therefore, the mean density of the block is as follows:

$$\rho_e = \rho_m - (\rho_m - \rho_f)\varnothing_0 + 0.5 \sum k_q q_1 \quad (q = x, y, z) \quad (22)$$

Isostatic equilibrium demands are as follows:

$$\rho_a H = z_1 \rho_e \quad (23)$$

For a constant rate of sedimentation  $dz/dt = h_s > 0$ , pressure difference per unit time that must drive the lithospheric block down:

$$d(\Delta P) / dt = g \{h_s \rho_e - H \rho_a\} \quad (24)$$

The rate of change of the force driving the block to sink is the following:

$$\begin{aligned} d(F_{\text{down}}) / dt &= d(\Delta P) / dt \times x_1 y_1 \\ &= g x_1 y_1 \{h_s \rho_e - H \rho_a\} \end{aligned} \quad (25)$$

“ $x_1 y_1$ ” being the horizontal basal area of the block on which the force works.

Dividing this by the mass of the block, as in Eq. 21, the rate of increase of acceleration of the sinking block is as follows:

$$d(f_{\text{down}}) / dt = d(F_{\text{down}}) / dt \prod q_1^{-1} \rho_e^{-1} \quad (q = x, y, z) \quad (26)$$

$$d(f_{\text{down}}) / dt = g z_1^{-1} \{h_s \rho_e - H \rho_a\} \{\rho_m - (\rho_m - \rho_f)\varnothing_0 + 0.5 \sum k_q q_1\}^{-1} \quad (q = x, y, z) \quad (27)$$

$$d(f_{\text{down}}) / dt = g z_1^{-1} \left[ h_s - H \rho_a \{\rho_m - (\rho_m - \rho_f)\varnothing_0 + 0.5 \sum k_q q_1\}^{-1} \right] \quad (q = x, y, z) \quad (28)$$

Consider contrarily that the lithospheric block is under a constant rate of erosion that decreases its height at a rate  $h_e = -dz/dt$ . In this case,

$$\begin{aligned} d(f_{\text{up}}) / dt \\ = g z_1^{-1} \left[ H \rho_a \{\rho_m - (\rho_m - \rho_f)\varnothing_0 + 0.5 \sum k_q q_1\}^{-1} - h_s \right] \end{aligned} \quad (29)$$

$\rho_f = 0$  can be put in Eqs. 21, 28, and 29 to find respective simpler expressions for fluid free “dry” blocks.

### Discussions and conclusions

Understanding mean density and isostasy helps to work out the geoid height anomaly, forward seismic modeling, and other geodetic/gravity issues (Fowler 2005). In such seismic modeling, Kirchhoff’s integral equation is encountered that involves pressure term (Benthien and Hobbs 2005). From the expression for mass, for case I, in Eq. 11, the pressure

exerted on the “ $xy$ ” area (see Fig. 1) is as follows:

Pressure  $P_I$

$$= \{\rho_m \prod q_1 - (\rho_m - \rho_f)\varnothing_0 b(e^{-bz_1} + 1)\} g x^{-1} y^{-1} \quad (30)$$

For case II,  $P_{II}$

$$= \prod q_1 \{\rho_m - (\rho_m - \rho_f)\varnothing_0 + 0.5 \sum k_q q_1\} x^{-1} y^{-1} \quad (31)$$

The second byproduct of this study is that one can deduce the Bouguer correction required for gravity station at the top (point M in Fig. 1) of the floating lithosphere. Following Eq. 5.35 of Fowler (2005), the Bouguer correction is as follows:

$$\delta g_B = 2 \pi G \rho_e' z_s \quad (32)$$

Here,  $G = 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$ : gravitational constant,  $z_s$ : height of the measuring point above the sea level, and  $\rho_e'$ : mean density of the material between the measurement point (M) and sea level.  $\rho_e'$  can be obtained by first finding the mass of the block RSVK in Fig. 1 as follows for both the cases I and II:

$$M_I = \int_0^{z_s} \int_0^{x_1} \int_0^{y_1} \rho_{\text{bw}}(x, y, z) \, dx \, dy \, dz \quad (33)$$

This would be followed by

$$\rho_e' = M_I g x^{-1} y^{-1} \quad (34)$$

Isostatic disequilibrium could develop by plate tectonic reasons (Stüwe 2007), or simply by either accumulation (here  $dz/dt > 0$ ) or removal (here  $dz/dt < 0$ ) of ice (McGuire 2012). So far geoscientists have deduced isostatic uplift rates (as distance per unit time) from a few places of the world, ranging from a fraction of a millimeter per year (Adams et al. 2010) up to several millimeter per year (Takano et al. 2012). The present work suggests that in geological time scale, however, isostatic disequilibrium should happen by speeding up uplift/subsidence of lithospheric blocks. Sometimes  $\rho_m$  and  $\rho_f$  are constant at different depths (Syvitski et al. 2007), and in other cases,  $\rho_f$  has been reported to increase with depth (Patwardhan 2012). Density of the fluid as well as the matrix must be ideally pressure and therefore depth dependent (Djomani et al. 2001). The presented model presumes depth independent  $\rho_m$  and  $\rho_f$  magnitudes. Attempting their depth dependence complicate the model (Appendix-I). For clastic rocks/sediments,  $\rho_m \sim 2.65 \text{ g cm}^{-3}$  (Cranganu 2009). Compressibility of the pore fluid or water is merely 0.44 (1/GPa). Its thermal expansion coefficient  $2 \times 10^{-4} \text{ (1/}^\circ\text{C)}$  is also low (Zhao et al. 2008). The “formation compressibility” of the bulk porous rocks ranges from  $3 \times 10^{-6}$  to  $25 \times 10^{-6}$  psi (Dandekar 2013), which is too miniscule to contribute significantly in

varying density with depth. In different considerations, the depth dependence of density ( $\rho_z$ ) can be expressed in ways other than Eq. 3. For example, Ray (2001) presented  $\rho_z$  as a function of density of crust at the surface, density and expansion coefficient of the mantle, thickness ratio between the thickened crust and the reference crust, and temperature at depth  $z$ . His equation did not consider bulk wet density ( $\rho_{bw}$ ), matrix density ( $\rho_m$ ), and fluid density ( $\rho_f$ ).

This work presumes constant density of asthenosphere (as in Xu et al. 2016 for the North China Craton). For regions with layered asthenosphere of different densities (e.g., Holtzman et al. 2003), the combined pressures exerted by all such layers of known thicknesses can however be easily accommodated in Eq. 9. On the other hand, if we comprehend that the asthenosphere layer is also characterized by smooth linear or exponential variation of density, its mean density can be estimated in the same way it has been done for the lithospheric block (Eq. 12).

Two more issues need to mention. First, sediments undergo compaction at some depth, beyond which obviously pore fluid in determining the mean density of that sediment/sedimentary rock is to be considered zero. Second, if a layer of (nearly) constant density, such as a salt bed, exists at the top or bottom or even inside the density-stratified lithospheric block, the presented equations must be modified. Such a correction would be required when more realistically the lithospheric block consists of a basement of crystalline rock at bottom and sediments/sedimentary rocks over it. The weight exerted by such a layer is to be added along with that of the remainder of the lithospheric slice. Dividing the total weight by the volume of the slice would yield the density (Eq. 12). Rise of buoyant hot magma at sub-surface would disturb the already established density structure considered in this work. The model then would require terrain-specific modification. Density of granite batholithos, possibly lying below sedimentary layers, can vary both vertically and laterally (Bott 1967), which could be systematic due to chemical changes and empirical equations can be set up (Oliver 1977).

Note that vertical brittle fault planes cannot be explained by Anderson’s or subsequent other theories of faulting (Misra and Mukherjee 2017). Isostatic disequilibrium (similar to Eq. 28 in the new derivation in this work) could be the favored model in such cases (e.g., Misra et al. 2014, 2015; Mukherjee 2014, 2015).

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### Appendix-I

The compressibility of the rock matrix, at constant temperature, is as follows:

$$\beta_1 = v_i^{-1} \Delta P (v_f - v_i) \tag{35}$$

Here,  $v_i$ : initial volume of the matrix,  $v_f$ : final volume of the matrix, and  $\Delta P$ : change in overburden pressure.

Or,

$$\rho_{mz} = \{ \rho_{m0} (1 + \Delta P \beta_1) \} \tag{36}$$

where  $\rho_{mz}$ : matrix density at depth  $z$  and  $\rho_{m0}$ : matrix density at surface.

Pressure at a depth “ $z$ ” by an overburden rock mass is

$$P = \rho_{ov} g z \tag{37}$$

$\rho_{ov}$ : Density of the overburden rocks/sediments.

Therefore,

$$\rho_{mz} = \{ \rho_{m0} (1 + \rho_{ov} g z \beta_1) \} \tag{38}$$

But  $\rho_{ov}$  also depends on  $\rho_m$  of the above rocks and eventually on  $\beta_1$ . This work seeks mean density (here  $\rho_{ov}$ ) so that the pressure exerted by the lithospheric block can be deduced. If we use such an expression of  $\rho_{mz}$ , the calculation remains unsolved. The same problem will hold true with  $\rho_{fz}$  (density of pore fluid at depth  $z$ ). A second point is that, compressibility of the matrix and that of the pore fluid at constant pressure, also needs attention. So Eq. 35 alone cannot explain the density of matrix (or the density of fluid from a similar equation) at depth  $z$ .

### Appendix-II: Symbols

- $\rho_{bw}$ : wet bulk density of sediments/rock
- $\rho_m$ : matrix/mineral/grain density
- $\rho_f$ : fluid density
- $\emptyset$ : fractional porosity
- $\emptyset_z$ : porosity at depth  $z$
- $\emptyset_0$ : porosity at surface
- $e$ : exponential series
- $b$ : a constant
- $\lambda$ : compaction constant
- $k_x, k_y, k_z$ : density gradients along X, Y, and Z directions, respectively
- $M$ : mass of lithospheric block
- $x_1, y_1, z_1$ : length, width, and thickness of lithospheric block
- $\rho_e$ : mean density of lithospheric block
- $\rho_e'$ : mean density of the material between the measurement point ( $M$ ) and sea level
- $\rho_a$ : density of asthenosphere
- $H$ : height of asthenosphere above compensation depth (see Fig. 1)

$h_s$ : sedimentation rate  
 $h_e$ : erosion rate  
 $\Delta P$ : pressure difference  
 $P_I$ : pressure in case I  
 $P_{II}$ : pressure in case II  
 $g$ : acceleration due to gravity  
 $F_{\text{down}}$ : downward acting force on lithospheric block  
 $f_{\text{down}}$ : downward acting acceleration on lithospheric block  
 $f_{\text{up}}$ : upward acting acceleration on lithospheric block  
 $\delta g_B$ : Bouguer gravity correction  
 $G$ : Gravitational constant  
 $z_s$ : height of the measuring point above the sea level  
 $\beta_1$ : compressibility of rock matrix  
 $v_i$ : initial volume of matrix  
 $v_f$ : final volume of matrix  
 $\rho_{mz}$ : matrix density at depth  $z$   
 $\rho_{m0}$ : matrix density at surface  
 $\rho_{ov}$ : density of overburden  
 $\rho_{fz}$ : density of pore fluid at depth  $z$

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